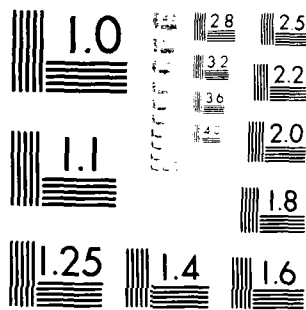


MODIFICATIONS TO ITERATIVE RECURSION UNFOLDING
ALGORITHMS AND COMPUTER CO..(U) NAVAL RESEARCH LAB
WASHINGTON DC K A LOWRY ET AL. 06 JUN 84 NRL-MR-5340

F/G 6/18

NL

END
DATE
FILMED
8-84
DTIC



MICROCOPY RESOLUTION TEST CHART
NATIONAL BUREAU OF STANDARDS-1963-A

AD-A142 475

2

NRL Memorandum Report 5340

**Modifications to Iterative Recursion Unfolding
Algorithms and Computer Codes to Find More
Appropriate Neutron Spectra**

K. A. LOWRY AND T. L. JOHNSON

Health Physics Staff

June 6, 1984



DTIC
ELECTE
JUN 27 1984
S B D

NAVAL RESEARCH LABORATORY
Washington, D.C.

Approved for public release; distribution unlimited.

84 06 26 097

DISC FILE COPY

REPORT DOCUMENTATION PAGE					
1a REPORT SECURITY CLASSIFICATION UNCLASSIFIED			1d RESTRICTIVE MARKINGS		
2a SECURITY CLASSIFICATION AUTHORITY			3 DISTRIBUTION AVAILABILITY OF REPORT		
2b DECLASSIFICATION DOWNGRADING SCHEDULE			Approved for public release; distribution unlimited		
4 PERFORMING ORGANIZATION REPORT NUMBER(S) NRL Memorandum Report 5340			5 MONITORING ORGANIZATION REPORT NUMBER(S)		
6a NAME OF PERFORMING ORGANIZATION Naval Research Laboratory		6b OFFICE SYMBOL (If applicable) Code 6070		7a NAME OF MONITORING ORGANIZATION	
6c ADDRESS (City, State and ZIP Code) Washington, DC 20375		7b ADDRESS (City, State and ZIP Code)			
8a NAME OF FUNDING SPONSORING ORGANIZATION Office of Naval Research		8b OFFICE SYMBOL (If applicable)		9 PROCUREMENT INSTRUMENT IDENTIFICATION NUMBER	
8c ADDRESS (City, State and ZIP Code) Arlington, VA 22217		10 SOURCE OF FUNDING NOS			
		PROGRAM ELEMENT NO 63571N		PROJECT NO S1580	TASK NO DN691-239 DN280-077
11 TITLE (Include Security Classification) (See page ii)					
12 PERSONAL AUTHOR(S) Lowry, K.A. and Johnson, T.L.					
13a TYPE OF REPORT Final		13b TIME COVERED FROM TO		14 DATE OF REPORT (Yr., Mo., Day) June 6, 1984	
15 PAGE COUNT 37					
16 SUPPLEMENTARY NOTATION					
17 COSATI CODES			18 SUBJECT TERMS (Continue on reverse if necessary and identify by block number)		
FIELD	GROUP	SUB GR			
			Neutrons Unfolding		
			Neutron spectrometry Computer code		
			(Continues)		
19 ABSTRACT (Continue on reverse if necessary and identify by block number) <p>➤ Modifications were made to iterative recursion unfolding algorithms and computer codes which allow the incorporation of prior knowledge, such as cutoff energy or preferred spectral shape, into the solution. Using these modifications, more appropriate neutron spectra are obtained.</p>					
20 DISTRIBUTION AVAILABILITY OF ABSTRACT UNCLASSIFIED UNLIMITED X SAME AS RPT DTIC USERS			21 ABSTRACT SECURITY CLASSIFICATION UNCLASSIFIED		
22a NAME OF RESPONSIBLE INDIVIDUAL T. L. Johnson			22b TELEPHONE NUMBER (Include Area Code) (202) 767-2232		22c OFFICE SYMBOL Code 6070

11. TITLE (Include Security Classification)

Modifications to Iterative Recursion Unfolding Algorithms and Computer Codes to Find More Appropriate Neutron Spectra

18. SUBJECT TERMS (Continued)

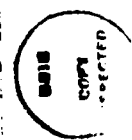
**Recursion unfolding
Bonner spheres
Linear equations
Radiation protection
Health physics
Neutron dosimetry**

CONTENTS

INTRODUCTION	1
EFFECT OF INITIAL SOLUTION AND SMOOTHING METHOD	2
BIASING THE SOLUTION TO OTHER STARTING SPECTRA	4
SMOOTHING TO OTHER STARTING SPECTRA	4
PROGRAM BUNKI	6
CONCLUSION	7
ACKNOWLEDGMENTS	7
REFERENCES	13
APPENDIX A — LISTING OF BUNKI COMPUTER CODE	15
APPENDIX B — DESCRIPTION OF INPUT PARAMETERS OF BUNKI	27
APPENDIX C — SAMPLE OF INTERACTIVE BUNKI SESSION	29
APPENDIX D — SAMPLE OUTPUTS OF BUNKI	31

DTIC
ELECTE
JUN 27 1984
B

✓	
A-1	



MODIFICATIONS TO ITERATIVE RECURSION UNFOLDING ALGORITHMS AND COMPUTER CODES TO FIND MORE APPROPRIATE NEUTRON SPECTRA

INTRODUCTION

The unfolding of neutron spectra using data from activation foils, Bonner spheres, or other detectors usually involves solving an equation of the form

$$Y_j = \int_{E_{\min}}^{E_{\max}} A_j(E) X(E) dE \quad (1)$$

$j = 1, 2, 3 \dots M$

where $X(E)$ represents the fluence distribution of the energy of the neutrons, Y_j is the response of the j th detector, $A_j(E)$ describes the response of the j th detector to neutrons of energy E , and M is the total number of detectors. Equation (1) is a degenerate case of a Fredholm integral equation of the first kind.

Although several methods exist for the formal solution of first-order integral equations, none of these methods are generally applicable when the detector response function $A_j(E)$ is not known analytically. This is the case for all practical systems used for neutron spectrometry. In practice, $A_j(E)$ is experimentally determined and/or calculated and is usually approximated by a response matrix having discrete values. Equation (1) is then replaced by a set of linear equations, and one is left with the problem of solving M equations in N unknowns, where M is the number of detectors available and N is the number of points needed to define the neutron spectrum. In matrix notation we have

$$\bar{Y} = A\bar{X} . \quad (2)$$

Usually N is greater than M , hence no unique solution exists, and even when $N \leq M$ the response matrix A is usually ill-conditioned resulting in wildly oscillating, sometimes negative, solutions having little physical significance. This led Gold [1] to introduce the terms "exact", "approximate", and

Manuscript approved March 2, 1984.

"appropriate" to characterize the solutions to equation (2). An exact solution, if one exists, satisfies the equation exactly but may be negative or oscillatory. Approximate solutions satisfy the equation only within reasonable error limits. From the approximate solutions, selection of the most physically acceptable solution yields an appropriate solution, which usually is not unique.

Methods for solving equations (1) and (2) for neutron spectrometry have been summarized by Nachtigall and Burger [2] and by Patterson and Thomas [3]. One method which finds a non-negative solution to equation (2) by minimizing through an iterative recursion procedure the deviation between the measured and calculated detector responses has been described by Scofield [4] and Gold [1]. This method was modified by O'Brien et al [5] and Sanna [6], who called the computer code used to unfold the spectrum BON31G. More recently, Dorashenko et al [7] have described another iterative recursion method for spectrum unfolding. A computer code SPUNIT, which uses this algorithm, has been written at Pacific Northwest Labs by Brackenbush and Scherpelz [8]. In addition to non-negativity of the solution, advantages usually listed for such iterative recursion methods of spectrum unfolding are ease of programming, and the possibility of adding smoothing to the solution. However, the application of other prior knowledge, such as cutoff energy or preferred spectral shape, has not been incorporated into the solution. It is our purpose to show how this may be accomplished in order to find more appropriate neutron spectra with these solution methods.

EFFECT OF INITIAL SOLUTION AND SMOOTHING METHOD

In Fig. 1 we show a typical neutron spectrum that has been unfolded from Bonner sphere data using BON31G or SPUNIT and the Sanna response matrix [9]. This spectrum was produced by moderating a Cf-252 fission source by placing it at the center of a 60 cm dia. steel sphere. The spectrum was further moderated by interposing a 60 cm x 60 cm x 3.8 cm thick Lucite slab between the steel ball and the Bonner sphere detectors. Detector counts were greater than 10K giving counting errors of approximately 1%. The spectrum shown in Fig. 1 is typical of spectra unfolded from Bonner sphere data in that there are few neutrons in the intermediate energy region (10^{-6} - 10^{-2} MeV) giving rise to the so-called "Bonner dip". Also, the high energy end of the spectrum

does not decrease as rapidly as calculations suggest [10]. Similarly shaped spectra are often also obtained with other unfolding codes such as LOUHI [11] which can incorporate prior knowledge into the solution [12, 13]. We shall see that the physically unreasonable spectrum shape shown in Fig. 1 is caused by the starting solution and the smoothing algorithm used.

It is usually stated that the shape of the initial solution has little or no influence on the final solution for recursion methods [6, 7]. That this is not the case is illustrated in Fig. 2. Using SPUNIT and the same data as in Fig. 1, spectra were unfolded using two drastically different initial solutions. One initial solution was a 1/E spectrum, i.e., a straight horizontal line on the lethargy scale shown. The second had alternate initial spectral values differing by a factor of 10. No smoothing was used in either case. Note that after 1000 iterations the spectrum unfolded with alternate factor of 10 initial values still retains this characteristic. Even after 5000 iterations this characteristic remained. Both spectra shown in Fig. 2 fit the experimental data with an average error less than 1%. The fact that both spectra fit the data so well indicates that Bonner sphere data cannot produce spectra having great resolution, hence their dependence on the initial solution. Using BON31G, we obtained similar results, however, 3000 iterations were necessary to fit the experimental data within 1%.

The smoothing algorithm also affects the shape of the spectrum shown in Fig. 1. The smoothing used was as follows:

$$X_s(i) = (SX(i-1) + X(i) + SX(i+1))/(1+2S), \quad i = 2, 3 \dots N-1 \quad (3)$$

$$X_s(1) = X(1), \quad (3a)$$

$$X_s(N) = X(N), \quad (3b)$$

where X is the value of spectrum before smoothing, X_s is the value of the spectrum after smoothing, and S is the smoothing factor. Smoothing was done on each iteration with $S = 0.01$. Sanna [6] used an equivalent equation with $S = 0.05$, smoothing on alternate iterations. $X(1)$ and $X(N)$ were also smoothed.

An examination of equation (3) reveals that this smoothing also tends to bias the solution to a 1/E spectrum. To eliminate this bias, Sanna later

removed smoothing completely [14, 15]. Hence, his earlier finding [6] that the final solution was independent of starting solution was probably caused by the introduction of smoothing. Using smoothing, we also obtained the same final solution for all starting spectra.

BIASING THE SOLUTION TO OTHER STARTING SPECTRA

The previous results indicate that the starting solution does affect the final solution, hence, it would seem appropriate to choose the most physically reasonable spectrum for a starting solution. Most neutron spectra encountered in radiation protection work are produced by the scattering, moderation, and absorption of neutrons originally produced by nuclear fission, particle accelerators, or by the α, n reaction from radioactive sources. These processes tend to produce spectra that can be characterized as having a high energy peak corresponding, or reduced somewhat by moderation, to the original neutron source energy, a $(1/E)^x$ intermediate energy component produced by elastic scattering, and a thermal peak whose magnitude is determined by the atomic number of the shielding and scattering material and by the thermal neutron absorption cross-section of these materials. We therefore developed an algorithm called MAXIET [16] for use in YOGI [17] to find the best solution made up of these three spectrum components. This spectrum can then be used as an initial solution. Using the same data as Fig. 1 and 2, the spectra obtained using only MAXIET, and using MAXIET for an initial solution for SPUNIT, are shown in Fig. 3. Using only MAXIET, the average error on the detector responses is 1.1%; using MAXIET plus SPUNIT this error is reduced to 0.5%. No smoothing was used in either case. Note that using MAXIET and SPUNIT gives a spectrum that fits the data better, and agrees better with calculated spectra [10], than is obtained using a $1/E$ initial solution.

SMOOTHING TO OTHER STARTING SPECTRA

As previously noted, equation (3) biases the solution to a straight line, i.e., $1/E$ spectrum. What is needed is some method to bias the solution toward any chosen spectrum, e.g., the MAXIET initial spectrum. Perhaps the easiest way to do this is to make a transformation on the response matrix and the

initial spectrum so that the transformed initial spectrum is a straight line, i.e.,

$$A^T(i,j) = A(i,j) X_I(i) \quad i = 1, 2 \dots N \quad (4)$$

$$j = 1, 2 \dots M$$

$$X_I^T(i) = 1 \quad i = 1, 2 \dots N \quad (5)$$

where A is the response matrix, A^T is the transformed response matrix, X_I is the initial spectrum, and X_I^T is the transformed initial spectrum.

The solution can then proceed using X_I^T and A^T to find the transformed solution X^T , using appropriate smoothing, e.g., equation (3). However, since the thermal neutron cross-sections of the shielding materials can tend to cause a very large or very small thermal neutron component, we do not smooth either the first or second energy intervals, i.e., equation (3) is replaced by

$$X_S^T(i) = (SX^T(i-1) + X^T(i) + SX^T(i+1))/(1+2S), \quad (6)$$

$$i = 3, 4 \dots N-1$$

$$X_S^T(1) = X^T(1) \quad (6a)$$

$$X_S^T(2) = X^T(2) \quad (6b)$$

$$X_S^T(N) = X^T(N) \quad (6c)$$

Alternatively, if one wishes to tie the Nth spectral value to a fixed (N+1)th value, this is easily accomplished by letting the index in equation (6) go to N and eliminating equation (6c).

After solving for X^T , the final solution is obtained by making the inverse transformation, i.e.,

$$X(i) = X_I(i) X^T(i), \quad i = 1, 2 \dots N, \quad (7)$$

To illustrate the use of smoothing to different starting spectra, errors of approximately 5% S.D. were introduced, using the algorithm in [6], to the

data used in the previous figures to simulate larger experimental errors. Spectra were then unfolded using a 1/E and a MAXIET starting spectrum using no smoothing, or equation (6) for smoothing with $S = 0.05$. These results are shown in Figs. 4 and 5. Note that using the MAXIET starting spectrum with smoothing to that spectrum gives more physically reasonable results than does the 1/E starting spectrum. More importantly, it can be seen that making the matrix transformation allows smoothing to any chosen initial spectrum.

PROGRAM BUNKI

The modifications which we have described to find more appropriate neutron spectra using BON31G or SPUNIT are incorporated in a computer program called BUNKI which is listed in Appendix A. BUNKI is written in FORTRAN IV and was coded for use on a DEC-10 computer. Tapes of the program and the response matrix file GIANT can be generated which are compatible with most mini or mainframe computers. Contact the authors for details.

The meanings of most of the variables used in BUNKI are apparent from their use in the program. The definitions of the parameters which affect the spectrum and which may be input by the user are listed in Appendix B. Those which are not usually changed are written directly into the program, but they could also be input by the user, for example, at lines 27400-28000. Numbers in parentheses refer to relevant line number(s) for the parameters.

BUNKI is written to be run interactively from a terminal. Appendix C is an example of a BUNKI session in which two spectra are generated using the same Bonner sphere input data. In the first case, the user inputs the initial spectrum; in the second, the MAXIET algorithm is used to find the initial spectrum. The unformatted variables written back to the terminal as a guide for the user are listed at lines 39200 and 59400 of BUNKI. The questions which BUNKI asks the user indicate the many options available such as: choice of detectors, choice of unfolding code, choice of response matrix, choice of initial spectrum, and choice of the number of energy intervals, in addition to the input parameters in Appendix B.

The results of the spectrum unfolding are written directly to disc in a file called FOR01.DAT which is automatically generated by the DEC-10. In

addition, the lethargy spectrum data may be written to a file called SPECX.DAT which is used in a separate program to create plots of the spectra. The output of the session listed in Appendix C is listed in Appendix D.

CONCLUSION

By properly choosing the initial solution, and by smoothing to that solution by means of a transformation on the detector response matrix, it is possible to incorporate prior knowledge into the neutron spectra obtained using iterative recursion unfolding algorithms. We have written a computer code called BUNKI which incorporates these modifications into BON31G and SPUNIT. These modifications should also be useful for use in other unfolding algorithms and computer codes.

ACKNOWLEDGMENTS

We thank Larry Brackenbush and Bob Scherpelz of Pacific Northwest Labs for bringing to our attention SPUNIT and the work of Doroshenko et al. / We are especially indebted to Bob Schwartz of the National Bureau of Standards whose question, "What causes that 'dip' in Bonner sphere spectra?", inspired us to do this work.

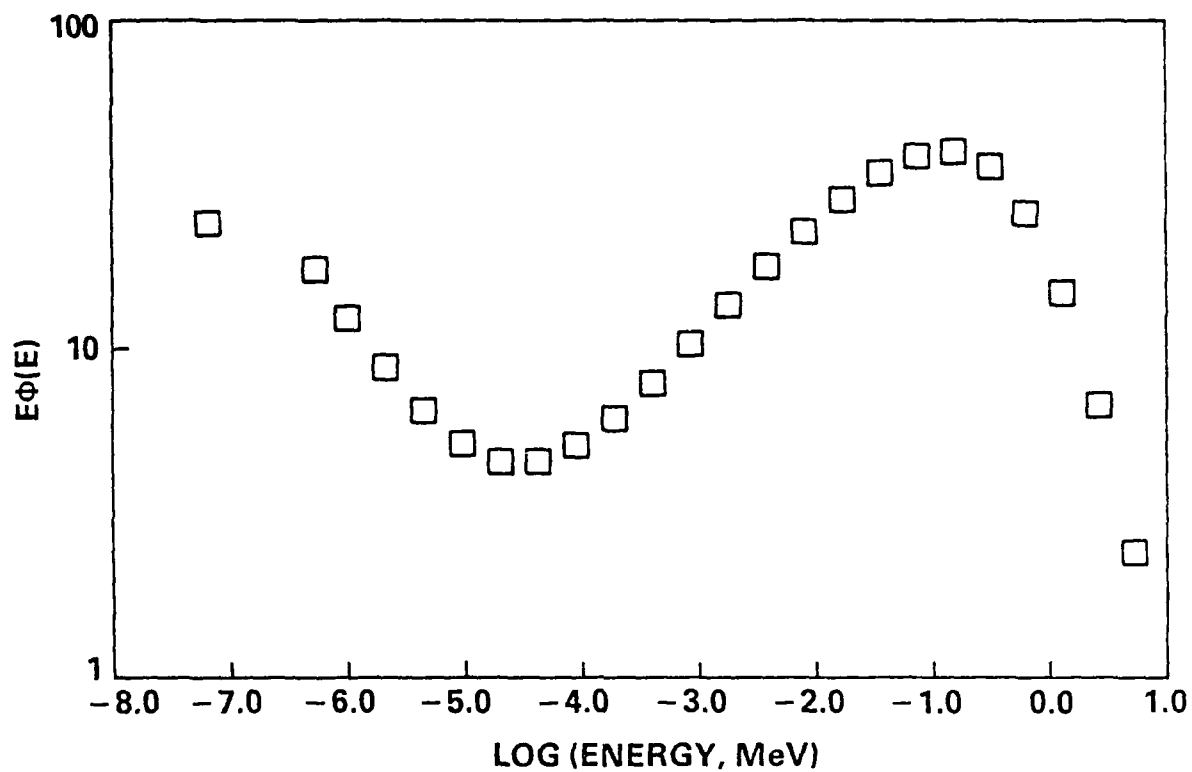


Fig. 1. Neutron spectrum of Cf-252 moderated by 30 cm of steel and 3.81 cm of Lucite. The spectrum was unfolded using SPUNIT (1000 iterations) from a $1/E$ starting spectrum and smoothed to that starting spectrum.

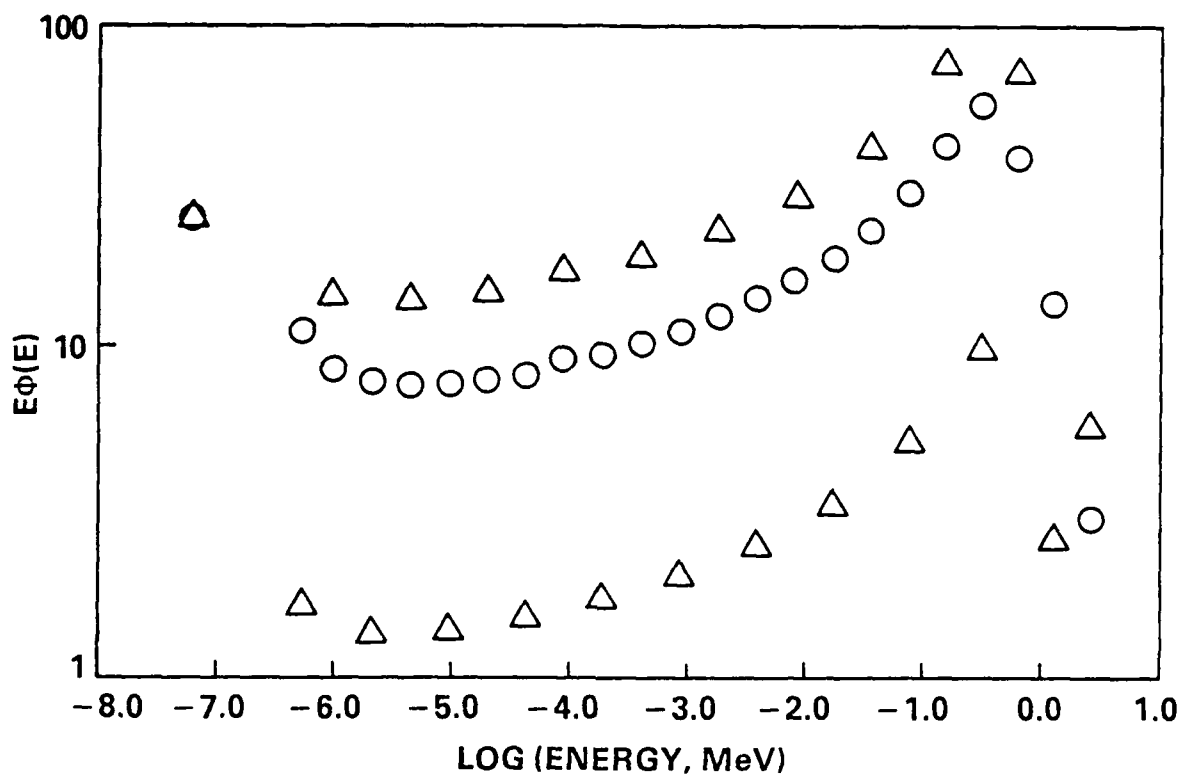


Fig. 2. Neutron spectra of Cf-252 moderated by 30 cm of steel and 3.81 cm of Lucite. The spectra were unfolded using SPUNIT (1000 iterations) using a $1/E$ starting spectrum (O), or using a starting spectrum with alternate initial values differing by a factor of 10 (Δ). No smoothing was used in either case.

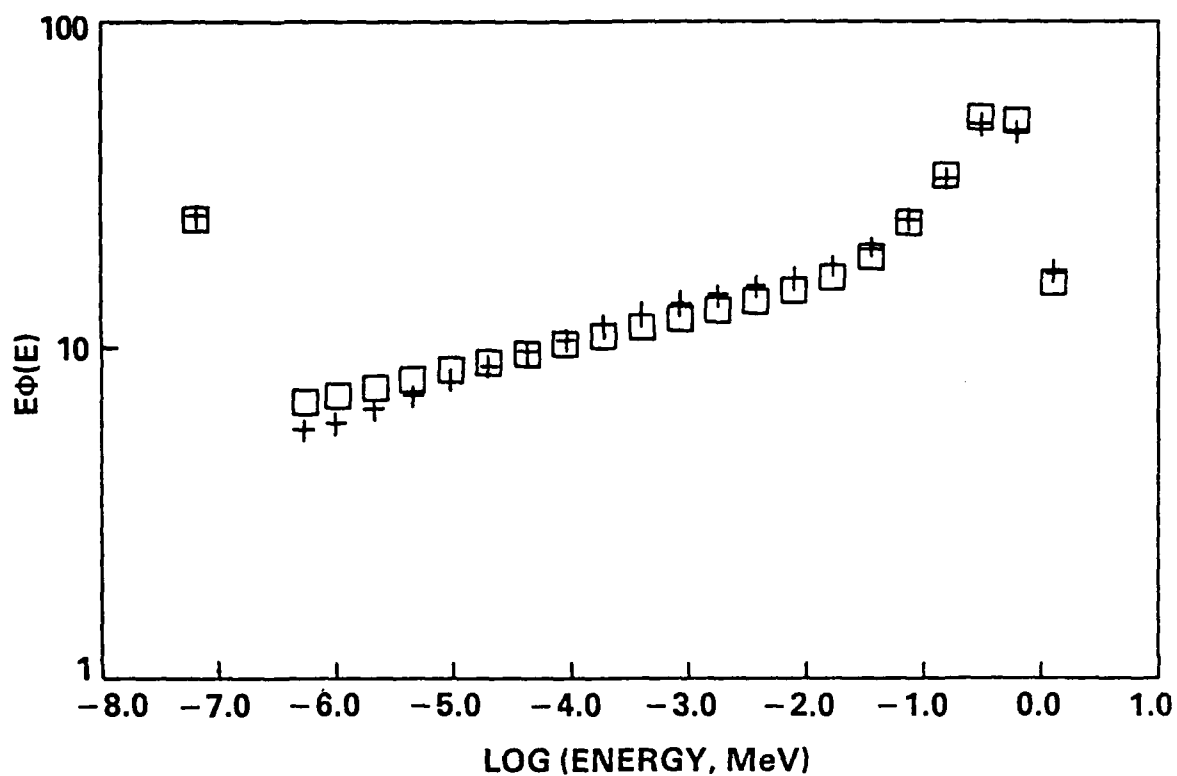


Fig. 3. Neutron spectra of Cf-252 moderated by 30 cm of steel and 3.81 cm of Lucite. The spectra were unfolded using the MAXIET algorithm (\square), and using the MAXIET spectrum as an initial solution for SPUNIT (1000 iterations) (+). No smoothing was used in either case.

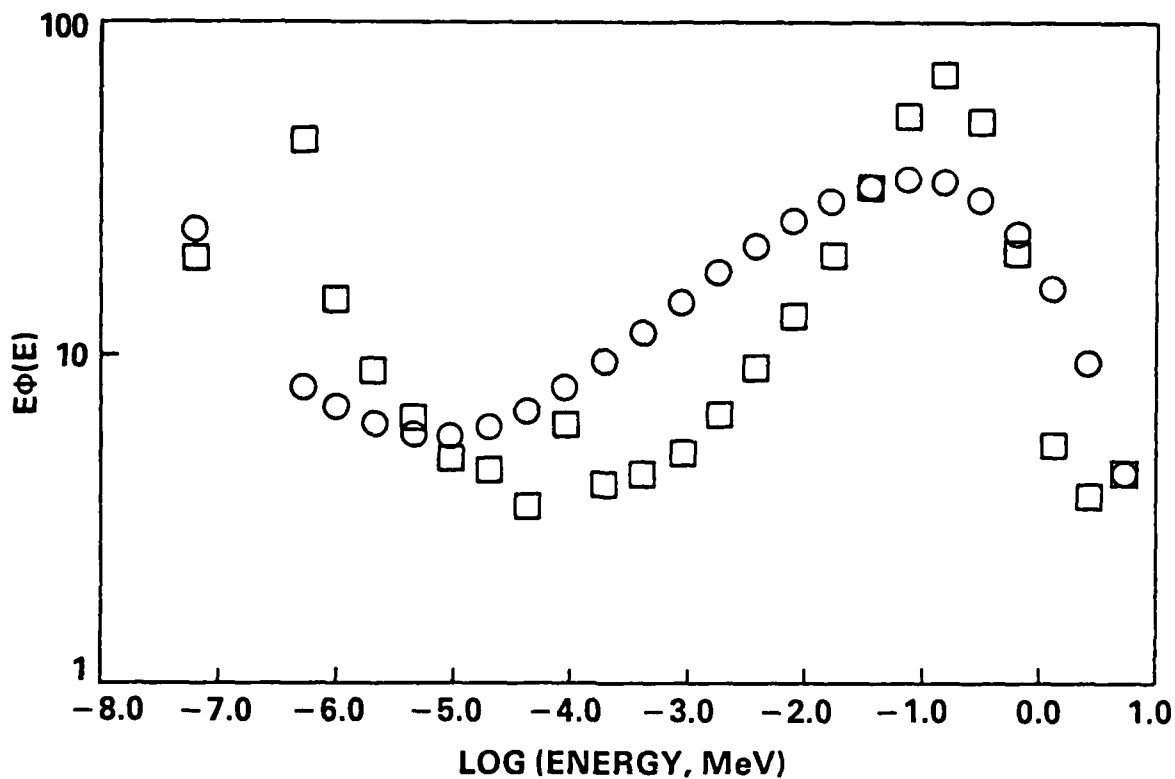


Fig. 4. Spectra unfolded using the same data as the previous figures with additional errors of 5% introduced to simulate larger experimental errors. Spectra were unfolded using SPUNIT (1000 iterations) from a $1/E$ starting spectrum with no smoothing (\square), or with smoothing to the starting spectrum (\circ), using the method described in the text.

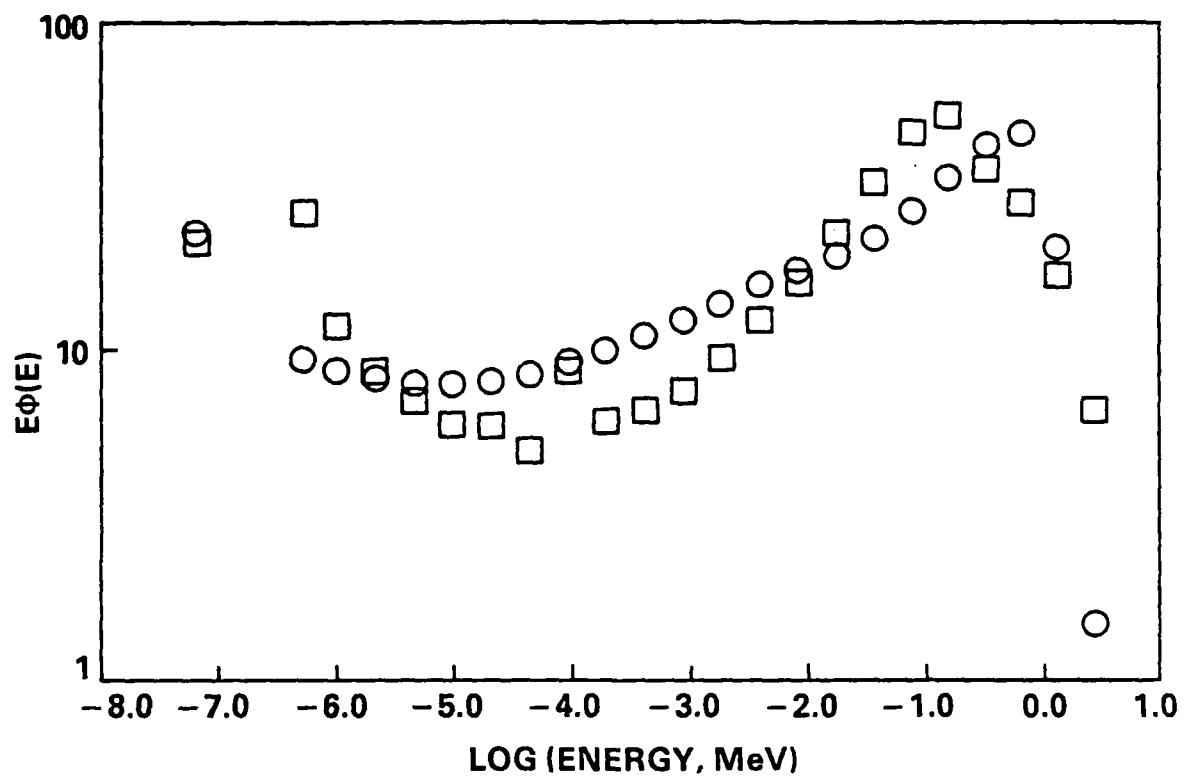


Fig. 5. Same as Fig. 4 except that MAXIET was used to determine the starting spectrum.

REFERENCES

1. Gold, R., "An Iterative Unfolding Method for Response Matrices", Argonne National Laboratory, Argonne, IL, ANL-6984 (1964).
2. Nachtigall, D. and Burger, G., "Dose Equivalent Determinations in Neutron Fields by Means of Moderator Techniques", in: Topics in Radiation Dosimetry, (Edited by F. H. Attix, W. C. Roesch and E. Tochlin) pp. 385-459 (New York: Academic Press) (1972).
3. Patterson, H. W., and Thomas, R. H., Accelerator Health Physics (New York: Academic Press) (1973).
4. Scofield, N. E., "A Technique for Unfolding Gamma-Ray Scintillation Spectrometer Pulse Height Distributions", U. S. Naval Radiological Defense Laboratory San Francisco, CA, USNRDL-TR-447 (1960).
5. O'Brien, K.; Sanna, R.; and McLaughlin, J.; "Inference of Accelerator Stray Neutron Spectra from Various Measurements", in: First Symposium on Accelerator Radiation Dosimetry and Experience, USAEC-CONF-651109 (1965).
6. Sanna, R. S., "Modification of an Iterative Code for Unfolding Neutron Spectra from Multisphere Data", Health and Safety Laboratory Energy Research and Development Administration New York, HASL-311 (1976).
7. Doroshenko, J. J.; Kraitov, S. N.; Kuznetsova, T.V.; Kushnereva, K. K.; and Leonov, E. S., "New Methods for Measuring Neutron Spectra with Energy from 0.4 eV to 10 MeV by Track and Activation Detectors", Nucl. Tech. 33, 296-304 (1977).
8. Brackenbush, L. W. and Scherpelz, R. I., "SPUNIT, A Computer Code for Multisphere Unfolding", in: Computer Applications in Health Physics, Proceedings of the Health Physics Topical Meeting, Pasco, WA (1984).
9. Sanna, R. S., "Thirty One Group Response Matrices for the Multisphere Neutron Spectrometer over the Energy Ranges Thermal to 400 MeV", Health and Safety Laboratory, U. S. Atomic Energy Commission, NY, HASL-767 (1973).
10. Ing, H. and Makra, S., Compendium of Neutron Spectra in Criticality Accident Dosimetry, International Atomic Energy Agency, Vienna, STI/DOC/10/180 (1978).
11. Routti, J. T. and Sandberg, J. V., "General Purpose Unfolding Program LOUHI78 with Linear and Nonlinear Regularizations", Comput. Phys. Commun. 21, 119-144, (1980).
12. Griffith, R. V.; Slaughter, D. R.; Patterson, H. W.; Beach, J. L.; Frank, E. G.; Rueppel, D. W.; and Fisher, J. C.; "Multi-Technique Characterization of Neutron Fields from Moderated ^{252}Cf and $^{238}\text{PuBe}$ Sources", in: Proceedings of National and International Standardization of Radiation Dosimetry. International Atomic Energy Agency, Vienna, IAEA STI/PUB/471 (1978).

13. Endres, G.W.R.; Aldrich, J. M.; Brackenbush, L. W.; Faust, L. G.; Griffith R. V.; and Hankins, D. E.; "Neutron Dosimetry at Commercial Nuclear Plants", Pacific Northwest Laboratory, Richland, WA, NUREG/CR-1769, PNL-3585 (1981).
14. Sanna, R.S.; Hajnal, F.; McLaughlin, J.E.; Gulbin, J. F.; and Ryan, R. M.; "Neutron Measurements Inside PWR Containments", Environmental Measurements Laboratory, U. S. Dept. of Energy, New York, EML-379 (1980).
15. Sanna, R. S., Private communication. Environmental Measurement Laboratory, New York (1983).
16. Johnson, T. L., "A Directed Search Algorithm for Finding Physically Reasonable Neutron Spectra from Bonner Sphere Data" to be presented to the Health Physics Society Annual Meeting, New Orleans, LA, June 1984.
17. Johnson, T. L. and Gorbics, S. G., "An Iterative Perturbation Method for Unfolding Neutron Spectra from Bonner Sphere Data", Health Phys. 41, 859 (1981).

APPENDIX A

LISTING OF BUNKI COMPUTER CODE

00100 C NOVEMBER 8,1983
00200
00300
00400 C BUNKI
00500
00600 C THIS PROGRAM CALCULATES THE PARTICLE FLUENCE, DOSE,
00700 C AND DOSE EQUIVALENT SPECTRA AS A FUNCTION OF NEUTRON
00800 C ENERGY USING THE BON3IG OR SPUNIT UNFOLDING CODES TO
00900 C SOLVE THE MATRIX APPROXIMATION OF A FREDHOLM INTEGRAL
01000 C EQUATION OF THE FIRST KIND. THE TOTAL FLUENCE, DOSE,
01100 C DOSE EQUIVALENT, QUALITY FACTOR, AVERAGE ENERGY, AND
01200 C RESPONSES OF SELECTED DETECTORS ARE ALSO CALCULATED.
01300 C THE INITIAL SPECTRUM MAY BE SPECIFIED BY THE USER OR
01400 C A 1/E, MAXWELLIAN, INITIAL SPECTRUM MAY BE DETERMINED
01500 C USING MAXIET, AN ALGORITHM ORIGINALLY DEVISED FOR USE
01600 C IN YOGI, THE ITERATIVE UNFOLDING CODE DEVELOPED AT THE
01700 C NAVAL RESEARCH LABORATORY. A MATRIX TRANSFORMATION
01800 C ALLOWS SMOOTHING BIASED TO EITHER INITIAL SPECTRUM.
01900 C
02000 C BUNKI WAS ORIGINALLY PROGRAMED AT THE NAVAL RESEARCH
02100 C LABORATORY IN JULY, 1983, BY KIMBERLY A. LOWRY AND
02200 C TOMMY L. JOHNSON.
02300 C
02400 C PROGRAM BUNKI
02500 C DOUBLE PRECISION BALL
02600 C DIMENSION ALETH(12,31),SPC(31),BCE(12),CTLD(31),
02700 C & BCC(12),CRAD(31),CREM(31),CNUTRK(31),HEAD(20),
02800 C & RAD(31),REM(31),EEND(32),CE(31),SPL(32),WDLETH(31),
02900 C & CNTA(31),CA70(31),CHAN(31),SPLPLT(31,10),SPLI(32),
03000 C & SPLMAX(32),LL(12),BALL(12),PCTERR(12),CODE(12),
03100 C & PREM(31),ERRBCE(12),WHTBCE(12),A(108,31),
03200 C & SPL(32),SS(32),BK(31,31),VECT(31)
03300 C ENERGY INTERVAL END POINTS
03400 C DATA EEND/
03500 C & 1.000E-08,4.140E-07,6.826E-07,1.445E-06,3.059E-06,
03600 C & 6.476E-06,1.371E-05,2.902E-05,6.144E-05,1.301E-04,
03700 C & 2.754E-04,5.929E-04,1.234E-03,2.613E-03,5.531E-03,
03800 C & 1.171E-02,2.479E-02,5.247E-02,1.111E-01,2.237E-01,
03900 C & 4.508E-01,9.072E-01,1.872E-00,3.679E-00,7.408E-00,
04000 C & 1.492E+01,2.581E+01,4.465E+01,7.725E+01,1.336E+02,
04100 C & 2.312E+02,4.000E+02/
04200 C NAMES OF THE DETECTORS
04300 C DATA BALL/
04400 C & 8HBARE ,8HBARE+CD ,8H2 INCH ,8H2"+CD ,
04500 C & 8H3 INCH ,8H3"+CD ,8H5 INCH ,8H5"+CD ,
04600 C & 8H8 INCH ,8H10 INCH ,8H12 INCH ,8H18 INCH /
04700 C FLUENCE TO DOSE CONVERSION FACTORS
04800 C DATA CRAD/
04900 C & 5.260E-10,6.088E-10,6.175E-10,6.135E-10,6.070E-10,
05000 C & 6.008E-10,5.970E-10,5.937E-10,5.892E-10,5.698E-10,
05100 C & 5.465E-10,5.251E-10,5.149E-10,5.083E-10,5.039E-10,
05200 C & 5.629E-10,6.639E-10,7.847E-10,1.038E-09,1.454E-09,
05300 C & 2.244E-09,3.406E-09,4.220E-09,5.778E-09,6.622E-09,
05400 C & 9.097E-09,9.610E-09,1.013E-08,1.137E-08,1.490E-08,
05500 C & 1.794E-08/
05600 C FLUENCE TO DOSE EQUIVALENT CONVERSION FACTORS

16

```

11800      MM=JJ
11900      C INPUT NUMBER AND KIND OF DETECTORS
12000      1030 IF(JX.EQ.1)GO TO 1060
12100      WRITE(5,1040)
12200      1040 FORMAT(' CHANGE DETECTORS?')
12300      1050 FORMAT(A1)
12400      READ(5,1050)CHDET
12500      IF(CHDET.EQ.'Y')GO TO 1060
12600      IF(CHDET.EQ.'N')GO TO 1140
12700      GO TO 1030
12800      1060 WRITE(5,1070)
12900      1070 FORMAT(' NUMBER OF DETECTORS?')
13000      READ(5,*)KK
13100      1080 WRITE(5,1090)
13200      1090 FORMAT(' TYPE DETECTOR CODES(? FOR HELP)')
13300      READ(5,1100)CODE(1)
13400      1100 FORMAT(A4)
13500      IF(CODE(1).EQ.'?')GO TO 1120
13600      READ(5,1100)(CODE(I),I=2,KK)
13700      DO 1110 I=1,KK
13800      LL(I)=0
13900      IF(CODE(I).EQ.'0')LL(I)=1
14000      IF(CODE(I).EQ.'0C')LL(I)=2
14100      IF(CODE(I).EQ.'2')LL(I)=3
14200      IF(CODE(I).EQ.'2C')LL(I)=4
14300      IF(CODE(I).EQ.'3')LL(I)=5
14400      IF(CODE(I).EQ.'3C')LL(I)=6
14500      IF(CODE(I).EQ.'5')LL(I)=7
14600      IF(CODE(I).EQ.'5C')LL(I)=8
14700      IF(CODE(I).EQ.'8')LL(I)=9
14800      IF(CODE(I).EQ.'10')LL(I)=10
14900      IF(CODE(I).EQ.'12')LL(I)=11
15000      IF(CODE(I).EQ.'18')LL(I)=12
15100      IF(LL(I).EQ.0)GO TO 1080
15200      1110 CONTINUE
15300      GO TO 1140
15400      1120 WRITE(5,1130)
15500      1130 FORMAT(10X,'0.....BARE',/,
15600      & 10X,'0C.....BARE+CADIUM COVER',/,
15700      & 10X,'2.....2 INCH BALL',/,
15800      & 10X,'2C.....2 INCH CADIUM COVERD BALL',/,
15900      & 10X,'3.....3 INCH BALL',/,
16000      & 10X,'3C.....3 INCH CADIUM COVERED BALL',/,
16100      & 10X,'5.....5 INCH BALL',/,
16200      & 10X,'5C.....5 INCH CADIUM COVERD BALL',/,
16300      & 10X,'8.....8 INCH BALL',/,
16400      & 10X,'10.....10 INCH BALL',/,
16500      & 10X,'12.....12 INCH BALL',/,
16600      & 10X,'18.....18 INCH BALL',/,/)
16700      GO TO 1080
16800      C SELECT RESPONSE MATRIX
16900      1140 WRITE(5,1145)
17000      1145 FORMAT(' TYPE MATRIX NAME(? FOR HELP)')
17100      READ(5,1100)RMTX
17200      IF(RMTX.EQ.'?')GO TO 1150
17300      LLL=1
17400      IF(RMTX.EQ.'SAN4')LLL=0
17500      IF(RMTX.EQ.'SAN1')LLL=12
17600      IF(RMTX.EQ.'M60')LLL=24
17700      IF(RMTX.EQ.'M65')LLL=36
17800      IF(RMTX.EQ.'MS13')LLL=48
17900      IF(RMTX.EQ.'LOGN')LLL=60

```

```

18000      IF(RMTX.EQ.'BARC')LLL=72
18100      IF(RMTX.EQ.'UTA4')LLL=84
18200      IF(RMTX.EQ.'UTA1')LLL=96
18300      IF(LLL.EQ.1)GO TO 1140
18400      GO TO 1160
18500      1150  WRITE(5,1155)
18600      1155  FORMAT(10X,'SANA.....SANA 4MM X 4MM ',/,
18700      & 10X,'SANA13.....SANA 13MM X 13MM ',/,
18800      & 10X,'M60.....M60',/,
18900      & 10X,'M65.....M65',/,
19000      & 10X,'MS13.....MODIFIED SANA 13MMX13MM',/,
19100      & 10X,'LOGNM.....LOGNORMAL ',/,
19200      & 10X,'BARC.....BHABA MONTE CARLO ',/,
19300      & 10X,'UTA4.....4MM X 4MM U OF TEXAS ',/,
19400      & 10X,'UTA13.....13MM X 13MM U OF TEXAS ',/)
19500      GO TO 1140
19600      C  SELECT UNFOLDING CODE
19700      1160  WRITE(5,1165)
19800      1165  FORMAT(' TYPE UNFOLDING CODE(? FOR HELP)')
19900      READ(5,1100)UNFOLD
20000      IF(UNFOLD.EQ.'?')GO TO 1170
20100      IF(UNFOLD.EQ.'BON3'.OR.UNFOLD.EQ.'SPUN')GO TO 1180
20200      GO TO 1160
20300      1170  WRITE(5,1175)
20400      1175  FORMAT(10X,'BON31G.....HASL UNFOLDING CODE',/,
20500      & 10X,'SPUNIT.....PNL-SOVIET UNFOLDING CODE',/)
20600      GO TO 1160
20700      1180  CONTINUE
20800      C  SELECT APPROPRIATE COLUMNS FROM MATRIX AND ADJUST
20900      C  MATRIX TO UNFOLD PER UNIT LETHARGY
21000      DO 1190 I=1,KK
21100      L=LL(I)+LLL
21200      DO 1190 J=1,JJJ
21300      1190  ALETH(I,J)=A(L,J)*WDLETH(J)
21400      C  INPUT VARIABLE INITIAL CONDITIONS AND DATA
21500      1200  WRITE(5,1210)
21600      1210  FORMAT(' PUT SPECTRA IN FILE FOR PLOTTING?')
21700      READ(5,1050)FILE
21800      1220  IF(FILE.EQ.'Y')GO TO 1230
21900      IF(FILE.EQ.'N')GO TO 1250
22000      GO TO 1200
22100      1230  WRITE(5,1240)
22200      1240  FORMAT(' FILE NUMBER OF THE SPECTRUM? (1 TO 10)')
22300      READ(5,*)KX
22400      IF(KX.LT.1.OR.KX.GT.10)GO TO 1230
22500      1250  IF(JX.EQ.1)GO TO 1270
22600      WRITE(5,1260)
22700      1260  FORMAT(' CHANGE THE HEADING?')
22800      READ(5,1050)CHHEAD
22900      IF(CHHEAD.EQ.'Y')GO TO 1270
23000      IF(CHHEAD.EQ.'N')GO TO 1320
23100      GO TO 1250
23200      1270  WRITE (5,1280)
23300      1280  FORMAT(' TYPE THE HEADING (<80 CHARACTERS)')
23400      READ(5,1290)(HEAD(I),I=1,20)
23500      1290  FORMAT(20A4)
23600      1320  WRITE(5,1330)
23700      1330  FORMAT(' SEARCH FOR MAXWELLIAN, 1/E INITIAL SPECTRUM?')
23800      READ(5,1050)SRHMA
23900      IF(SRHMA.NE.'Y'.AND.SRHMA.NE.'N')GO TO 1320

```

```

24000      IF(SRHMA.NE.SRHMAX)LX=1
24100      SRHMAX=SRHMA
24200      IF(SRHMAX.EQ.'Y')GO TO 1380
24300      IF(SRHMAX.EQ.'N')GO TO 1340
24400      GO TO 1320
24500 1340  IF(LX.EQ.1)GO TO 1360
24600      IF(CHNUM.EQ.'Y')GO TO 1360
24700      WRITE(5,1350)
24800 1350  FORMAT(' CHANGE INITIAL SPECTRUM?')
24900      READ(5,1050)CHISPC
25000      IF(CHISPC.EQ.'Y')GO TO 1360
25100      IF(CHISPC.EQ.'N')GO TO 1480
25200      GO TO 1340
25300 1360  WRITE(5,1370)JJJ
25400 1370  FORMAT(' TYPE INITIAL SPECTRUM (' ,I2,' VALUES')')
25500      READ(5,*)(SPLI(I),I=1,JJJ)
25600      GO TO 1480
25700 1380  IF(LX.EQ.1)GO TO 1400
25800      WRITE(5,1390)
25900 1390  FORMAT(' CHANGE MAXWELLIAN TEMP, SHAPE, OR PERTURBATION?')
26000      READ(5,1050)CHTEMP
26100      IF(CHTEMP.EQ.'Y')GO TO 1400
26200      IF(CHTEMP.EQ.'N')GO TO 1480
26300      GO TO 1380
26400 1400  WRITE(5,1410)
26500 1410  FORMAT(' TYPE MAXWELLIAN TEMP, SHAPE, AND PERTURBATION')
26600      READ(5,*)TEMPIJ,SHAPE,PERTMP
26700 1480  IF(JX.EQ.1)GO TO 1500
26800      WRITE(5,1490)
26900 1490  FORMAT(' CHANGE FIT PARAMETERS?')
27000      READ(5,1050)CHFIT
27100      IF(CHFIT.EQ.'Y')GO TO 1500
27200      IF(CHFIT.EQ.'N')GO TO 1520
27300      GO TO 1480
27400 1500  WRITE(5,1510)
27500 1510  FORMAT(' TYPE: END TEST(%),',
27600      & /,' SMOOTHING FACTOR, ',
27700      & /,' CALIBRATION FACTOR, ',
27800      & /,' ITERATIONS BEFORE ERROR TEST, ',
27900      & /,' AND MAXIMUM NUMBER OF ITERATIONS')
28000      READ(5,*)TSTPER,SMO,CAL,ITRTST,ITRMAX
28100      PERTHM=1.0+20*PERSLP
28200      PERE=1.0+10*PERSLP
28300      IF(SRHMAX.EQ.'N'.AND.ITRMAX.EQ.0)GO TO 1820
28400      TSTPER=(KK*TSTPER**2)/10000
28500 1520  TEMPI=TEMPIJ
28600      SLOPEI=SLOPEJ
28700      THERMI=THERMJ
28800      IF(JX.EQ.1)GO TO 1540
28900      IF(CHDET.EQ.'Y')GO TO 1540
29000 1525  WRITE(5,1530)
29100 1530  FORMAT(' CHANGE BALL DATA?')
29200      READ(5,1050)CHBCE
29300      IF(CHBCE.EQ.'Y')GO TO 1540
29400      IF(CHBCE.EQ.'N')GO TO 1590
29500      GO TO 1525
29600 1540  DO 1560 I=1,KK
29700      L=LL(I)
29800      WRITE(5,1550)BALL(L)
29900 1550  FORMAT(' TYPE ',A8,'BONNER SPHERE COUNT, % ERROR')

```

```

30000 1560 READ(5,*)(BCE(I),ERRBCE(I))
30100 C SUM ERRORS, MAKE DEAD TIME CORRECTION TO BALL COUNTS
30200 SUMWHT=0
30300 DO 1570 I=1, KK
30400 SUMWHT=SUMWHT+ERRBCE(I)
30500 1570 BCE(I)=BCE(I)/(1.0-BCE(I)*DEAD)
30600 C CALCULATE BALL COUNT ERROR WEIGHTS
30700 DO 1580 I=1, KK
30800 1580 WHTBCE(I)=SUMWHT/(KK*ERRBCE(I))
30900 C CALCULATE MAXWELLIAN, 1/E SPECTRUM IF REQUIRED
31000 1590 IF(SRMX.EQ.'N')GO TO 1820
31100 C BEGINNING OF MAXIET ALGORITHM
31200 C INITIALIZE FIT PARAMETERS
31300 ERROR=123456789
31400 TEMP=TEMP1
31500 ERRORM=ERROR
31600 1600 SLOPE=SLOPE1
31700 THERM=THERM1
31800 C CALCULATE MAXWELLIAN SPECTRUM
31900 SPMX=0
32000 DO 1610 I=1, JJJ
32100 SPLMAX(I)=(CE(I)**1.5)*(EXP(-CE(I)/TEMP))
32200 IF(SPLMAX(I).GT.SPMX)SPMX=SPLMAX(I)
32300 IF(SPMX.EQ.SPLMAX(I))GO TO 1610
32400 SPLMAX(I)=SPMX**SHAPE*SPLMAX(I)**(1.0-SHAPE)
32500 IF(SPLMAX(I).LT.SPLMAX(I-1)*SHP)SPLMAX(I)=SPLMAX(I-1)*SHP
32600 1610 CONTINUE
32700 C CALCULATE 1/E INITIAL SPECTRUM
32800 ERROR=123456789
32900 ERROR=123456789
33000 HGTE=SPMX*PERE
33100 1620 ERROR=ERROR
33200 1630 HGTE=HGTE/PERE
33300 1640 DO 1650 I=1, JJJ
33400 1650 SPLI(I)=HGTE*CE(I)**SLOPE
33500 C COMBINE MAXWELLIAN AND 1/E SPECTRA
33600 DO 1660 I=1, JJJ
33700 IF(SPLI(I).LT.SPLMAX(I))GO TO 1670
33800 SPLI(I)=(SPLI(I)+SPLMAX(I))*0.5
33900 1660 CONTINUE
34000 GO TO 1630
34100 1670 DO 1680 J=1, JJJ
34200 1680 SPLI(J)=SPLMAX(J)
34300 C ADJUST THERMAL ENERGY BIN
34400 SPLI(1)=SPLI(2)*THERM
34500 C CALCULATE SPHERE RESPONSES AND SUM FROM SPECTRUM
34600 DO 1690 M=1, KK
34700 BCC(M)=0
34800 DO 1690 J=1, JJJ
34900 1690 BCC(M)=BCC(M)+ALETH(M,J)*SPLI(J)
35000 C CALCULATE SUMS OF SPHERE DATA
35100 SUMBCE=0
35200 SUMBCC=0
35300 DO 1700 I=1, KK
35400 SUMBCE=SUMBCE+BCE(I)
35500 SUMBCC=SUMBCC+BCC(I)
35600 C NORMALIZE CALCULATED SPHERE RESPONSES
35700 C TO EXPERIMENTAL DATA
35800 RNORM=SUMBCE/SUMBCC
35900 DO 1710 I=1, KK

```

```

36000 1710 BCC(I)=BCC(I)*RNORM
36100 C CALCULATE ERROR ON FIT
36200 ERROR=0
36300 DO 1720 I=1,KK
36400 ERR=(BCC(I)-BCE(I))/BCE(I)
36500 1720 ERROR=ERROR+WHTBCE(I)*ERR*ERR
36600 IF(ERROR.LT.ERRORT)GO TO 1620
36700 HGTE=HGTE*PERE
36800 1730 IF(ERRORT.GE.ERRORRE)GO TO 1740
36900 C SAVE BEST VALUES OF FIT PARAMETERS
37000 ERRORRE=ERRORT
37100 HGTEE=HGTE
37200 THERME=THERM
37300 SLOPEE=SLOPE
37400 MX=0
37500 C CHANGE SLOPE
37600 1740 SLOPE=SLOPE+PERSLP
37700 IF(MX.EQ.1)SLOPE=SLOPEE
37800 C CHANGE THERMAL BIN
37900 THERM=THERME*PERTHM
38000 IF(THERM.GE.THMMAX)GO TO 1750
38100 IF(MX.EQ.0)THERM=THERME
38200 MX=MX+1
38300 IF(MX.GT.10)GO TO 1750
38400 HGTE=HGTE*PERE*(1.0+10*PERSLP)
38500 IF(MX.EQ.1)HGTE=HGTEE*PERE*PERTHM
38600 C RESET ERROR, SEARCH FOR BETTER FIT PARAMETERS
38700 ERRORT=123456789
38800 GO TO 1640
38900 C CALCULATE ERROR ON FIT
39000 1750 PERROR=100*(ERRORRE/KK)**.5
39100 C WRITE BEST VALUES OF FIT PARAMETERS TO TERMINAL
39200 WRITE(5,*)TEMP,SHAPE,HGTEE,SLOPEE,THERME,PERROR
39300 IF(ERRORRE.GE.ERRORRM)GO TO 1760
39400 C SAVE BEST VALUES OF FIT PARAMETERS
39500 TEMP=TEMP
39600 HGTEM=HGTEE
39700 THERNM=THERME
39800 SLOPEM=SLOPEE
39900 ERRORM=ERRORRE
40000 C CHANGE MAXWELLIAN TEMPERATURE IF REQUIRED
40100 IF(PERTMP.EQ.0)GO TO 1760
40200 TEMP=TEMP-PERTMP
40300 C RETURN AND SEARCH FOR BETTER PARAMETERS IF
40400 C MAXWELLIAN TEMP IS IN RANGE
40500 IF(TEMP.GT.PERTMP)GO TO 1600
40600 1760 CONTINUE
40700 C IF FINAL PARAMETERS EQUAL INITIAL PARAMETERS CHANGE
40800 C INITIAL PARAMETERS AND CONTINUE SEARCH
40900 A1=0
41000 IF(SLOPEM.EQ.SLOPEI)A1=1
41100 A2=0
41200 IF(THERNM.EQ.THERMI.AND.THERMI.GE.THMMIN)A2=1
41300 A3=0
41400 IF(TEMPM.EQ.TEMPI.AND.TEMPM.LT.TEMPIJ+10*PERTMP)A3=1
41500 IF(PERTMP.EQ.0)A3=0
41600 IF(A1.EQ.1)SLOPEI=SLOPEI-10*PERSLP
41700 IF(A2.EQ.1)THERMI=THERMI/PERTHM**3
41800 IF(A3.EQ.1)TEMPI=TEMPI+3*PERTMP
41900 IF(A1+A2+A3.GT..5)GO TO 1590

```

```

42000 C IF NOT, CALCULATE INITIAL SPECTRUM WITH BEST PARAMETERS
42100 SPMX=0
42200 DO 1770 I=1, JJJ
42300 SPLMAX(I)=(CE(I)**1.5)*(EXP(-CE(I)/TEMPM))
42400 IF(SPLMAX(I).GT.SPMX)SPMX=SPLMAX(I)
42500 IF(SPMX.EQ.SPLMAX(I))GO TO 1770
42600 SPLMAX(I)=SPMX**SHAPE*SPLMAX(I)**(1.0-SHAPE)
42700 IF(SPLMAX(I).LT.SPLMAX(I-1)*SHP)SPLMAX(I)=SPLMAX(I-1)*SHP
42800 1770 CONTINUE
42900 DO 1780 I=1, JJJ
43000 1780 SPLI(I)=HGTEM*CE(I)**SLOPEM
43100 DO 1790 I=1, JJJ
43200 IF(SPLI(I).LT.SPLMAX(I))GO TO 1800
43300 SPLI(I)=(SPLI(I)+SPLMAX(I))*0.5
43400 1790 CONTINUE
43500 1800 DO 1810 J=I, JJJ
43600 1810 SPLI(J)=SPLMAX(J)
43700 SPLI(1)=SPLI(2)*THERMM
43800 C COMPLETION OF MAXIET ALGORITHM
43900 C TRANSFORM MATRIX TO CONSTANT INITIAL SPECTRUM
44000 1820 DO 1825 I=1, JJJ
44100 DO 1825 K=1, KK
44200 1825 ALETH(K, I)=ALETH(K, I)*SPLI(I)
44300 DO 1830 I=1, JJJ
44400 1830 SPL(I)=1
44500 C CALCULATE SPHERE RESPONSES AND SUM FROM INITIAL SPECTRUM
44600 DO 1840 M=1, KK
44700 BCC(M)=0
44800 DO 1840 J=1, JJ
44900 1840 BCC(M)=BCC(M)+ALETH(M, J)*SPL(J)
45000 IF(SRHMAX.EQ.'N'.AND.ITRMAX.EQ.0)GO TO 2000
45100 C NORMALIZE CALCULATED SPHERE RESPONSES AND INITIAL
45200 C SPECTRUM TO EXPERIMENTAL DATA
45300 SUMBCC=0
45400 SUMBCE=0
45500 DO 1850 I=1, KK
45600 SUMBCE=SUMBCE+BCE(I)
45700 1850 SUMBCC=SUMBCC+BCC(I)
45800 RNORM=SUMBCE/SUMBCC
45900 DO 1860 I=1, KK
46000 1860 BCC(I)=BCC(I)*RNORM
46100 SUMBCC=SUMBCE
46200 DO 1870 I=1, JJJ
46300 SPL(I)=SPL(I)*RNORM
46400 1870 CONTINUE
46500 C CALCULATE ERROR ON FIT
46600 ERROR=0
46700 DO 1880 I=1, KK
46800 ERR=(BCC(I)-BCE(I))/BCE(I)
46900 1880 ERROR=ERROR+WHTECE(I)*ERR*ERR
47000 IF(ITRMAX.EQ.0)GO TO 2000
47100 ITER=0
47200 1890 ERRORU=ERROR
47300 C SELECT UNFOLDING ALGORITHM
47400 IF(UNFOLD.EQ.'BON3')GO TO 1945
47500 C BEGINNING OF SPUNIT UNFOLDING ALGORITHM
47600 IF(ITER.GT.0)GO TO 1905
47700 DO 1900 J=1, JJ
47800 SS(J)=0
47900 DO 1900 I=1, KK

```

```

48000 1900 SS(J)=SS(J)+ALETH(I,J)/BCE(I)
48100 1905 DO 1940 K=1,ITRTST
48200 ITER=ITER+1
48300 DO 1910 J=1,JJ
48400 SPLL(J)=0
48500 DO 1910 I=1,KK
48600 1910 SPLL(J)=SPLL(J)+(SPL(J)*ALETH(I,J))/(SS(J)*BCC(I,))
48700 DO 1920 J=3,MM
48800 1920 SPL(J)=(SPL(J-1)*SMO+SPLL(J)+SPLL(J+1)*SMO)/(1+2*SMO)
48900 SPL(1)=SPLL(1)
49000 SPL(2)=SPLL(2)
49100 IF(MM.EQ.JJ)GO TO 1925
49200 SPL(JJ)=SPLL(JJ)
49300 1925 DO 1930 M=1,KK
49400 BCC(M)=0
49500 DO 1930 J=1,JJ
49600 1930 BCC(M)=BCC(M)+ALETH(M,J)*SPL(J)
49700 1940 CONTINUE
49800 C END OF SPUNIT UNFOLDING ALGORITHM
49900 GO TO 1990
50000 C BEGINNING OF BON31G UNFOLDING ALGORITHM
50100 1945 IF(ITER.GT.0)GO TO 1960
50200 DO 1950 I=1,JJ
50300 DO 1950 J=1,JJ
50400 BK(J,I)=0
50500 DO 1950 M=1,KK
50600 1950 BK(J,I)=BK(J,I)+ALETH(M,J)*ALETH(M,I)
50700 DO 1955 I=1,JJ
50800 VECT(I)=0
50900 DO 1955 J=1,KK
51000 1955 VECT(I)=VECT(I)+ALETH(J,I)*BCE(J)
51100 1960 DO 1980 N=1,ITRTST
51200 ITER=ITER+1
51300 DO 1970 J=1,JJ
51400 AX=0.
51500 DO 1965 M=1,JJ
51600 1965 AX=SPL(M)*BK(J,M)+AX
51700 1970 SPLL(J)=SPL(J)*VECT(J)/AX
51800 DO 1975 J=3,MM
51900 1975 SPL(J)=(SPLL(J-1)*SMO+SPLL(J)+SPLL(J+1)*SMO)/(1+2*SMO)
52000 SPL(1)=SPLL(1)
52100 SPL(2)=SPLL(2)
52200 IF(MM.EQ.JJ)GO TO 1980
52300 SPL(JJ)=SPLL(JJ)
52400 1980 CONTINUE
52500 DO 1985 M=1,KK
52600 BCC(M)=0
52700 DO 1985 J=1,JJ
52800 1985 BCC(M)=BCC(M)+ALETH(M,J)*SPL(J)
52900 C END OF BON31G UNFOLDING ALGORITHM
53000 1990 CONTINUE
53100 C CALCULATE ERROR ON FIT
53200 ERROR=0
53300 DO 1995 I=1,KK
53400 ERR=(BCC(I)-BCE(I))/BCE(I)
53500 1995 ERROR=ERROR+WHTECE(I)*ERR*ERR
53600 C TEST FOR COMPLETION, CONTINUE IF NOT COMPLETE
53700 IF(ERROR/ERRORU.LT.TSTRAT.AND.ITER+ITRTST.LE.ITRMAX
53800 & .AND.ERROR.GT.TSTPER)GO TO 1890
53900 C IF COMPLETE, DO INVERSE TRANSFORM OF SPECTRUM AND MATRIX

```

```

54000 2000 DO 2005 I=1,JJJ
54100 2005 SPL(I)=SPL(I)*SPLI(I)
54200 DO 2010 I=1,KK
54300 DO 2010 J=1,JJJ
54400 2010 ALETH(I,J)=ALETH(I,J)/SPLI(J)
54500 C CALCULATE OUTPUT VALUES
54600 IF(SRHMAX.EQ.'N'.AND.ITRMAX.EQ.0)GO TO 2030
54700 2015 HGTEM=0.5*HGTEM/SPMX
54800 SUMERR=0
54900 DO 2020 I=1,KK
55000 PCTERR(I)=100*(BCC(I)-BCE(I))/BCE(I)
55100 2020 SUMERR=SUMERR+PCTERR(I)*PCTERR(I)
55200 PERROR=(SUMERR/KK)**.5
55300 2030 SUMSPC=0
55400 SUMRAD=0
55500 SUMREM=0
55600 SUMEXS=0
55700 SUMTLD=0
55800 SUMHAN=0
55900 SUMNTR=0
56000 SUMNTA=0
56100 SUMA70=0
56200 DO 2035 I=1,31
56300 IF(I.GT.JJ)SPL(I)=0
56400 SPL(I)=SPL(I)*CAL
56500 SPLPLT(I,KK)=SPL(I)
56600 SPC(I)=SPL(I)*WDLETH(I)
56700 SUMSPC=SUMSPC+SPC(I)
56800 REM(I)=CREM(I)*SPC(I)
56900 SUMREM=SUMREM+REM(I)
57000 RAD(I)=CRAD(I)*SPC(I)
57100 SUMRAD=SUMRAD+RAD(I)
57200 SUMEXS=SUMEXS+CE(I)*SPC(I)
57300 SUMTLD=SUMTLD+CTLD(I)*SPC(I)
57400 SUMHAN=SUMHAN+CHAN(I)*SPC(I)
57500 SUMNTR=SUMNTR+CNUTRK(I)*REM(I)
57600 SUMNTA=SUMNTA+CNNTA(I)*REM(I)
57700 2035 SUMA70=SUMA70+CA70(I)*SPC(I)
57800 QF=SUMREM/SUMRAD
57900 AVEEN=(SUMEXS-CE(1)*SPC(1))/(SUMSPC-SPC(1))
58000 SUMTLD=(SUMTLD/SUMREM)/4.064E+06
58100 SUMHAN=(SUMHAN/SUMREM)/2.085E+06
58200 SUMNTR=(SUMNTR/SUMREM)/.5748
58300 SUMNTA=(SUMNTA/SUMREM)/8.033
58400 SUMA70=(SUMA70/SUMREM)/4.755E+06
58500 DO 2040 I=1,JJ
58600 2040 PREM(I)=100*(REM(I)/SUMREM)
58700 C SET UNCOMPUTED PARAMETERS TO 0, WRITE PARAMETERS TO TERMINAL
58800 IF(SRHMAX.EQ.'N')THERMM=0
58900 IF(SRHMAX.EQ.'N')TEMPM=0
59000 IF(SRHMAX.EQ.'N')HGTEM=0
59100 IF(SRHMAX.EQ.'N')SLOPEM=0
59200 IF(SRHMAX.EQ.'N'.AND.ITRMAX.EQ.0)PERROR=0
59300 IF(ITRMAX.EQ.0)ITER=0
59400 WRITE(5,*)TEMPM,SHAPE,HGTEM,SLOPEM,THERMM,PERROR,ITER
59500 C DECIDE IF RESULTS ARE WORTH KEEPING
59600 2045 WRITE(5,2050)
59700 2050 FORMAT(' SAVE THESE RESULTS?')
59800 READ(5,1050)SAVE
59900 IF(SAVE.EQ.'N')GO TO 2190

```

```

60000      IF(SAVE.EQ.'Y')GO TO 2055
60100      GO TO 2045
60200      C WRITE RESULTS TO DISC
60300      2055  WRITE(1,2060)(HEAD(I),I=1,20)
60400      2060  FORMAT('1',20A4)
60500      IF(ITRMAX.EQ.0)GO TO 2080
60600      WRITE(1,2070)RMTX,UNFOLD,TEMPM,SHAPE,CAL,SNO,PERROR,ITER
60700      2070  FORMAT(1X,/,4X,'RESPONSE UNFOLD MAXWELL CALIB.
60800      &      SMOOTH PER CENT NO. OF',/,
60900      &      4X,'MATRIX CODE TEMP,SHAPE FACTOR FACTOR ERROR
61000      &      ITERATIONS',
61100      &      /,1X,A8,A10,FS.2,',',F4.2,F8.4,2(F10.4),I8)
61200      GO TO 2100
61300      2080  WRITE(1,2090) RMTX,TEMPM,SHAPE,HGTEM,SLOPEM,THERMM,PERROR,CAL
61400      2090  FORMAT(1X,/,4X,'RESPONSE MAXWELL 1/E X LETH
61500      &      THERMAL PERCENT CALIB.',/,
61600      &      4X,'MATRIX TEMP,SHAPE FACTOR SLOPE FACTOR ERROR
61700      &      FACTOR',
61800      &      /,1X,A8,F9.2,',',F4.2,F8.4,4(F10.4))
61900      2100  WRITE(1,2110)
62000      2110  FORMAT(1X/10X,9HDETECTORS,10X,8HMEASURED,5X,
62100      &      12H CALCULATED,5X,7HPERCENT,/,
62200      &      30X,6HCOUNTS,10X,6HCOUNTS,6X,10HDIFFERENCE)
62300      DO 2120 I=1,KK
62400      L=LL(I)
62500      2120  WRITE(1,2130),BALL(L),BCE(I),BCC(I),PCTERR(I)
62600      2130  FORMAT(10X,A10,1X,OPF15.3,OPF16.3,OPF15.3)
62700      WRITE(1,2140) SUMSPC,AVEEN,SUMRAD,SUMREM,QF,SUNTLD,SUMHAN,
62800      &      SUMNTR,SUMNTA,SUMA70
62900      2140  FORMAT(/,
63000      &      10X,'TOTAL FLUENCE=',T34,1PE11.3,T48,'NEUTRONS/CM2',/,
63100      &      10X,'AVE ENERGY(LESS TH)=' ,T34,1PE11.3,T48,'MEV',/,
63200      &      10X,'DOSE=' ,T34,1PE11.3,T48,'RAD',/,
63300      &      10X,'DOSE EQUIVALENT=' ,T34,1PE11.3,T48,'REM',/,
63400      &      10X,'QUALITY FACTOR=' ,T34,OPF7.3,T48,'REM/RAD',/,
63500      &      10X,'NRL TLD RESPONSE=' ,T34,F7.3,T48,'REM/REM(CF-252)',/,
63600      &      10X,'HANKINS TLD RESPONSE=' ,T34,F7.3,T48,'REM/REM(CF-252)',/,
63700      &      10X,'NEUTRAK RESPONSE=' ,T34,F7.3,T48,'REM/REM(CF-252)',/,
63800      &      10X,'NTA RESPONSE=' ,T34,F7.3,T48,'REM/REM(CF-252)',/,
63900      &      10X,'ANPDR-70 RESPONSE=' ,T34,F7.3,T48,'REM/REM(CF-252)',/)
64000      WRITE(1,2150)
64100      2150  FORMAT(3X,
64200      &      ' BIN ENERGY FLUENCE FLUENCE
64300      &      DOSE DOSE EQV. DOSE EQV.',/,
64400      &      ' NO. MAX (MEV) NEUT/CM2
64500      &      N/CM2/LETH (RAD) (REM) (% OF TOTAL)')
64600      DO 2160 I=1,JJ
64700      2160  WRITE(1,2170) I,EEND(I+1),SPC(I),SPL(I),RAD(I),
64800      &      REM(I),PREM(I)
64900      2170  FORMAT(2X,I4,2X,6(1PE11.3))
65000      WRITE(1,2180)
65100      2180  FORMAT(1X,/,/)
65200      2190  CONTINUE
65300      C SET FLAGS, RETURN FOR ANOTHER SPECTRUM IF DESIRED
65400      JX=0
65500      LX=0
65600      CHDET='N'
65700      CHNUM='N'
65800      2200  WRITE(5,2210)
65900      2210  FORMAT(' LAST SPECTRUM?')

```

```

66000      READ(5,1050)LASTSP
66100      IF(LASTSP.EQ.'N')GO TO 2220
66200      IF(LASTSP.EQ.'Y')GO TO 2240
66300      GO TO 2200
66400 2220  WRITE(5,2230)
66500 2230  FORMAT(' CHANGE MATRIX, DETECTORS, UNFOLDING CODE',/,
66600      &  ' AND/OR ENERGY INTERVALS?')
66700      READ(5,1050)CHMTX
66800      IF(CHMTX.EQ.'N')GO TO 1220
66900      IF(CHMTX.EQ.'Y')GO TO 1010
67000      GO TO 2220
67100 2240  IF(FILE.EQ.'N')GO TO 2250
67200 C  WRITE SPECTRA TO FILE FOR PLOTTING IF REQUIRED
67300      OPEN(UNIT=1,FILE='SPECX')
67400      WRITE(1,*)SPLPLT
67500      CLOSE(UNIT=1,FILE='SPECX')
67600 2250  END

```

APPENDIX B

Description of Input Parameters of BUNKI

- SLOPEJ (9800,38600) The initial value of the slope of the 1/E part of the MAXIET spectrum. Usually 0.
- PERSLP (9900,37600,28600,38400,41600) The amount by which the slope of the 1/E part of the MAXIET spectrum is changed in searching for a better fit to the data. Typical values, 0.005 - 0.02.
- THERMJ (10000,28700) The initial value of the thermal bin of the initial MAXIET spectrum. Usually 1.0.
- THMMIN (10100,41200) The minimum value allowable for the thermal bin of the MAXIET spectrum.
- THMMAX (10200,38000) The maximum allowable value for the thermal bin of the MAXIET spectrum. THMMAX and THMMIN are set from physical characteristics of the radiation environment.
- DEAD (10300,30500) The dead time of the instrument used to determine the detector counts.
- SHP (10400,32500,42700) The minimum value of the (I+1) bin relative to the I bin for the initial MAXIET spectrum. Used to limit the high energy roll-off of the calculated Maxwellian spectrum.
- TSTRAT (10500,53700) The maximum allowable value of the error on the fit relative to the value when the error was last tested. Prevents further iterations when no significant improvement in fit is occurring with further iterations. If set > 1.0 this test will not terminate the fit. Typical values: 0.9, 0.99, 0.999, 0.9999, 1.1.
- TEMPIJ (25900-26600,28500,41400) The initial guess of the temperature of the Maxwellian peak. Program asks for "Maxwellian Temp.".
- SHAPE (25900-26600, 36400,42600) The shape of the high temperature portion of the Maxwellian peak. Typical values, 0-0.5. Originally the program searched for the best shape, but this feature was considered unnecessary and is now user input. Program asks for "shape".
- PERTMP (25900-26600, 40100,40200,40500,41400,41500,41800) The amount by which the Maxwellian temperature is changed in searching for a better fit to the data. Should be approximately 10% the Maxwellian temperature, TEMPIJ. May be positive, negative, or zero. Positive searches for a lower temperature, negative for a higher temperature. If set to zero, the fit is forced from TEMPIJ which is often a useful feature. Program asks for "perturbation".
- ISTPER (26900-28000,28400,53800) An end test used to terminate the fit. When the error on the fit drops below this error the fit is terminated. Program asks for "end test (%)".

APPENDIX B (Con't)

- SMO (26900-28000, 48800,51900) The smoothing factor. Typical values, 0-0.05. Smoothing increases with increasing SMO. Program asks for "smoothing factor".
- CAL (26900-28000, 56400,60600) A calibration factor used to correct the spectrum to agree with some calibration standard. Typical values we have found for a 4 mm x 4 mm crystal and the Sanna matrix are 1.2-1.6 for Cf-252. Program asks for "calibration factor".
- ITRTST (26900-28000, 48100,51100,53700) The number of iterations before making a test to decide if the fit should be terminated. Typical values, 1-100. Program asks for "iterations before error test".
- ITRMAX (26900-28000, 45000,47000,53700,54600,59200,59300,60500) The maximum number of iterations allowed. Typically 100-1000. May also be set to 0. If the user inputs the initial spectrum and ITRMAX=0 the program does not ask for detector data but calculates the output parameters directly from the input spectrum. This option is useful for calculating integral parameters from known spectra. If using the MAXIET algorithm, setting ITRMAX=0 prevents any further fitting of the data by either SPUNIT or BON31G. The MAXIET spectrum, and calculated integral parameters, are then output in a slightly different format than that shown in Appendix D. Program asks for "maximum number of iterations".
- PERTHM (28100,37900,38500,41700) The amount by which the thermal bin of the MAXIET spectrum is changed in searching for a better fit to the data. PERTHM is not input directly, but is calculated from PERSLP.
- PERE (28200,33000,33200,36700,38400,38500) The amount by which the magnitude of the 1/E part of the MAXIET spectrum is changed in searching for a better fit to the data. PERE is not input directly, but is calculated from PERSLP. A reasonable relationship between PERSLP, PERTHM, and PERE must be maintained to insure that the MAXIET algorithm will find the best fit to the detector data.

APPENDIX C SAMPLE OF INTERACTIVE BUNKI SESSION

.LOGIN 720,1564
 JOB 32 RNL-603A-DUAL-KI-54 TTY127
 [LGNJSP Other jobs same PPN:63]
 [LGNRDU Recomputing disk usage]
 [LGNQTA DSKU IN:3000 OUT:3000 USED:1605]
 1510 17-Feb-84 Fri
 [No mail]

.EX BUNKI.FOR
 LINK: Loading
 [LNKXCT BUNKI execution]

NUMBER OF ENERGY INTERVALS?
25
 NUMBER OF DETECTORS?
7
 TYPE DETECTOR CODES (? FOR HELP)

0
2C
3C
5
8
10
12

TYPE MATRIX NAME (? FOR HELP)
SAN4

TYPE UNFOLDING CODE (? FOR HELP)
SPUNIT

PUT SPECTRA IN FILE FOR PLOTTING?
N

TYPE THE HEADING (<80 CHARACTERS)
CF 252 in 60 CM STEEL BALL + 1.27 CM LUCITE

SEARCH FOR MAXWELLIAN, 1/E INITIAL SPECTRUM?

N
 TYPE INITIAL SPECTRUM(26 VALUES)

26*1
 TYPE: END TEST(%),
 SMOOTHING FACTOR,
 CALIBRATION FACTOR,
 ITERATIONS BEFORE ERROR TEST,
 AND MAXIMUM NUMBER OF ITERATIONS

.2,0,1,25,1000

TYPE BARE BONNER SPHERE COUNT, % ERROR

935,3

TYPE 2"+CD BONNER SPHERE COUNT, % ERROR

2125,1

TYPE 3"+CD BONNER SPHERE COUNT, % ERROR

5197,1

TYPE 5 INCH BONNER SPHERE COUNT, % ERROR

8361,1

TYPE 8 INCH BONNER SPHERE COUNT, % ERROR

4547,1

TYPE 10 INCH BONNER SPHERE COUNT, % ERROR

2305,1

TYPE 12 INCH BONNER SPHERE COUNT, % ERROR

1063,1

0.0000000E+00, 0.0000000E+00, 0.0000000E+00, 0.0000000E+00,
0.0000000E+00, 0.4940264, 1000,

SAVE THESE RESULTS?

Y

LAST SPECTRUM?

N

CHANGE MATRIX, DETECTORS, UNFOLDING CODE
AND/OR ENERGY INTERVALS?

N

CHANGE THE HEADING?

N

SEARCH FOR MAXWELLIAN, 1/E INITIAL SPECTRUM?

Y

TYPE MAXWELLIAN TEMP, SHAPE, AND PERTURBATION

.5,0,.05

CHANGE FIT PARAMETERS?

Y

TYPE: END TEST(%)
SMOOTHING FACTOR,
CALIBRATION FACTOR,
ITERATIONS BEFORE ERROR TEST,
AND MAXIMUM NUMBER OF ITERATIONS

1,0,1,1,1000

CHANGE BALL DATA?

N

0.5000000	,	0.0000000E+00,	0.1458820	,	0.1400000	,
2.985984	,	10.29025	,			
0.4500000	,	0.0000000E+00,	0.1255103	,	0.1500000	,
2.985984	,	8.285000	,			
0.4000000	,	0.0000000E+00,	0.1118169	,	0.1800000	,
4.299817	,	5.430347	,			
0.3500000	,	0.0000000E+00,	0.8898502E-01	,	0.1900000	,
5.159780	,	3.348522	,			
0.3000000	,	0.0000000E+00,	0.6416974E-01	,	0.2100000	,
6.191737	,	1.054216	,			
0.2500000	,	0.0000000E+00,	0.9536391E-02	,	0.5000000E-01,	
2.488320	,	1.539098	,			
0.3000000	,	0.0000000E+00,	0.5167460	,	0.2100000	,
6.191737	,	0.7344218	,	1,		

SAVE THESE RESULTS?

Y

LAST SPECTRUM?

Y

END OF EXECUTION

CPU TIME: 33.33 ELAPSED TIME: 3:54.08

EXIT

APPENDIX D SAMPLE OUTPUTS OF BUNKI

(specification): b:steve2
 Viewing file 'b:steve2'. Hit <space> to continue; 'Alt-V' to terminate.

TYPE BUNN.DAT
 1 CF 252 IN 60 CM STEEL BALL + 1.27 CM LUCITE

RESPONSE MATRIX	UNFOLD CODE	MAXWELL TEMP, SHAPE	CALIB. FACTOR	SMOOTH FACTOR	PER CENT ERROR	NO. OF ITERATIONS
SAN4	SPUN	0.00, .00	1.0000	0.0000	0.4940	1000

DETECTOR	MEASURED COUNTS	CALCULATED COUNTS	PERCENT DIFFERENCE
BARE	935.000	935.163	0.017
2"CD	2125.000	2121.371	-0.171
3"CD	5197.000	5223.262	0.505
5-INCH	8361.000	8307.583	-0.639
8 INCH	4547.000	4579.329	0.711
10 INCH	2305.000	2289.550	-0.670
12 INCH	1063.000	1065.617	0.246

TOTAL FLUENCE=	3.426E+04	NEUTRONS/CM2
AVE ENERGY(LESS TH)=	3.104E-01	MEV
DOSE=	4.115E-05	RAD
DOSE EQUIVALENT=	3.372E-04	REM
QUALITY FACTOR=	8.194	REM/RAD
NRL TLD RESPONSE=	6.570	REM/REM(CF-252)
"HANKINS" TLD RESPONSE=	9.521	REM/REM(CF-252)
NEUTRAK RESPONSE=	0.521	REM/REM(CF-252)
NTA RESPONSE=	0.400	REM/REM(CF-252)
ANPDR-70 RESPONSE=	0.992	REM/REM(CF-252)

BIN NO.	ENERGY MAX (MEV)	FLUENCE NEUT/CM2	FLUENCE N/CM2/LETH	DOSE (RAD)	DOSE EQV. (REM)	DOSE EQV. (% OF TOTAL)
1	4.140E-07	4.409E+03	2.727E+03	2.319E-06	5.075E-06	1.505E+00
2	6.826E-07	3.949E+02	1.818E+03	2.404E-07	4.865E-07	1.443E-01
3	1.445E-06	3.489E+02	1.071E+03	2.154E-07	4.375E-07	1.298E-01
4	3.059E-06	2.968E+02	9.112E+02	1.821E-07	3.692E-07	1.095E-01
5	6.476E-06	2.814E+02	8.640E+02	1.708E-07	3.453E-07	1.024E-01
6	1.371E-05	2.778E+02	8.529E+02	1.669E-07	3.364E-07	9.978E-02
7	2.902E-05	2.939E+02	9.024E+02	1.755E-07	3.512E-07	1.042E-01
8	6.144E-05	3.167E+02	9.722E+02	1.880E-07	3.734E-07	1.107E-01
9	1.301E-04	4.070E+02	1.249E+03	2.398E-07	4.726E-07	1.402E-01
10	2.754E-04	3.889E+02	1.194E+03	2.216E-07	4.368E-07	1.295E-01
11	5.929E-04	4.741E+02	1.424E+03	2.591E-07	5.121E-07	1.519E-01
12	1.234E-03	5.177E+02	1.626E+03	2.718E-07	5.384E-07	1.597E-01
13	2.613E-03	6.231E+02	1.912E+03	3.208E-07	6.356E-07	1.885E-01
14	5.531E-03	7.779E+02	2.389E+03	3.954E-07	7.842E-07	2.326E-01
15	1.171E-02	9.992E+02	3.067E+03	5.035E-07	1.008E-06	2.990E-01
16	2.479E-02	1.324E+03	4.066E+03	7.456E-07	2.005E-06	5.947E-01
17	5.247E-02	1.870E+03	5.741E+03	1.241E-06	5.023E-06	1.490E+00
18	1.111E-01	2.903E+03	8.909E+03	2.278E-06	1.386E-05	4.111E+00
19	2.237E-01	4.483E+03	1.475E+04	4.654E-06	3.721E-05	1.104E+01
20	4.508E-01	6.445E+03	2.118E+04	9.371E-06	9.126E-05	2.707E+01
21	9.072E-01	4.617E+03	1.520E+04	1.036E-05	1.104E-04	3.274E+01
22	1.872E+00	1.422E+03	4.520E+03	4.844E-06	4.986E-05	1.479E+01
23	3.679E+00	2.886E+02	9.834E+02	1.218E-06	1.153E-05	3.419E+00
24	7.408E+00	7.049E+01	2.319E+02	4.073E-07	2.871E-06	8.513E-01
25	1.492E+01	2.413E+01	7.935E+01	1.598E-07	9.919E-07	2.942E-01

1 CF 252 IN 60 CM STEEL BALL + 1.27 CM LUCITE

RESPONSE MATRIX	UNFOLD CODE	MAXWELL TEMP, SHAPE	CALIB. FACTOR	SMOOTH FACTOR	PER CENT ERROR	NO. OF ITERATIONS
SAN4	SPUN	0.30, .00	1.0000	0.0000	0.7344	1

DETECTOR	MEASURED COUNTS	CALCULATED COUNTS	PERCENT DIFFERENCE
BARE	935.000	931.928	-0.329
2"+CD	2125.000	2105.156	-0.934
3"+CD	5197.000	5234.559	0.723
5 INCH	8361.000	8283.147	-0.931
8 INCH	4547.000	4596.901	1.097
10 INCH	2305.000	2303.379	-0.070
12 INCH	1063.000	1067.717	0.444

TOTAL FLUENCE=	3.420E+04	NEUTRONS/CM2
AVE ENERGY(LESS TH)=	2.948E-01	MEV
DOSE=	4.106E-05	RAD
DOSE EQUIVALENT=	3.396E-04	REM
QUALITY FACTOR=	8.270	REM/RAD
NRL TLD RESPONSE=	6.597	REM/REM(CF-252)
"HANKINS" TLD RESPONSE=	8.929	REM/REM(CF-252)
NEUTRAK RESPONSE=	0.492	REM/REM(CF-252)
NTA RESPONSE=	0.399	REM/REM(CF-252)
ANPDR-70 RESPONSE=	0.999	REM/REM(CF-252)

BIN NO.	ENERGY MAX (MEV)	FLUENCE NEUT/CM2	FLUENCE N/CM2/LETH	DOSE (RAD)	DOSE EQV. (REM)	DOSE EQV. (% OF TOTAL)
1	4.140E-07	4.709E+03	2.912E+03	2.477E-06	5.420E-06	1.596E+00
2	6.826E-07	1.015E+02	4.675E+02	6.180E-08	1.251E-07	3.683E-02
3	1.445E-06	1.732E+02	5.317E+02	1.069E-07	2.172E-07	6.394E-02
4	3.059E-06	2.024E+02	6.215E+02	1.242E-07	2.518E-07	7.415E-02
5	6.476E-06	2.367E+02	7.267E+02	1.437E-07	2.904E-07	8.552E-02
6	1.371E-05	2.768E+02	8.498E+02	1.663E-07	3.352E-07	9.871E-02
7	2.902E-05	3.236E+02	9.938E+02	1.932E-07	3.868E-07	1.139E-01
8	6.144E-05	3.787E+02	1.162E+03	2.248E-07	4.464E-07	1.315E-01
9	1.301E-04	4.430E+02	1.360E+03	2.610E-07	5.144E-07	1.515E-01
10	2.754E-04	5.181E+02	1.591E+03	2.952E-07	5.819E-07	1.713E-01
11	5.929E-04	6.211E+02	1.865E+03	3.394E-07	6.708E-07	1.975E-01
12	1.234E-03	6.953E+02	2.184E+03	3.651E-07	7.232E-07	2.129E-01
13	2.613E-03	8.335E+02	2.558E+03	4.292E-07	8.501E-07	2.503E-01
14	5.531E-03	9.817E+02	3.014E+03	4.990E-07	9.895E-07	2.914E-01
15	1.171E-02	1.170E+03	3.592E+03	5.895E-07	1.180E-06	3.476E-01
16	2.479E-02	1.431E+03	4.393E+03	8.055E-07	2.167E-06	6.380E-01
17	5.247E-02	1.850E+03	5.690E+03	1.228E-06	4.970E-06	1.463E+00
18	1.111E-01	2.615E+03	8.025E+03	2.052E-06	1.249E-05	3.676E+00
19	2.237E-01	3.653E+03	1.202E+04	3.792E-06	3.032E-05	8.927E+00
20	4.508E-01	5.633E+03	1.851E+04	8.191E-06	7.977E-05	2.349E+01
21	9.072E-01	5.486E+03	1.806E+04	1.231E-05	1.312E-04	3.862E+01
22	1.872E+00	1.807E+03	5.745E+03	6.156E-06	6.337E-05	1.866E+01
23	3.679E+00	5.889E+01	2.007E+02	2.485E-07	2.352E-06	6.927E-01
24	7.408E+00	6.098E-01	2.006E+00	3.524E-09	2.483E-08	7.312E-03
25	1.492E+01	6.099E-03	2.006E-02	4.039E-11	2.507E-10	7.383E-05

